# CN excitation in cold molecular clouds and in circumstellar envelope

# dépasser les frontières MWWW.cnrrs.fr

Yulia Kalugina<sup>1,2</sup>, François Lique<sup>1</sup>, Ghassen Saidani<sup>3</sup>, Aline Gardez<sup>3</sup>, Ludovic Biennier<sup>3</sup> and Robert Georges<sup>3</sup>

 <sup>1</sup>LOMC – UMR 6294, CNRS-Université du Havre, 25 rue Philippe Lebon, BP 540, 76058 Le Havre, France
 <sup>2</sup>Department of Optics and Spectroscopy, Tomsk State University, 36 Lenin Avenue, Tomsk 634050, Russia
 <sup>3</sup>UMR 6251, CNRS-Université de Rennes 1, Campus de Beaulieu 35042 Rennes Cedex, France



## INTRODUCTION

•CN molecule is one of the most widely distributed in ISM.
•The CN thermal lines probe dense regions in molecular clouds and circulstellar envelopes.

RC+10216

•Modelling of molecular emission spectra from ISM requires the knowledge of accurate rate coefficients for excitation by collisions with the most abundant species. Without these rates, only approximate estimates of the molecular

At low temperatures, collisional energy transfer affects the population of rotational levels of CN. We present state-to-state rate coefficients between the first 25 fine and 76 hyperfine levels of  $CN(X^2\Sigma^+)$ , being in the ground vibrational state, in collisions with  $H_2(j = 0)$ . The new fine and hyperfine-structure resolved rate coefficients are the first ones obtained for the CN radical in collision with  $H_2$ . In circumstellar envelope, CN radical can be produced by photodissociation of HCN. The products could be both in the ground and excited vibrational states. Hovewer, CN is observed only in its ground vibrational state. Thus, it is crucial to study the vibrational de-excitation of CN by the most abundant collisional partners. We present the study of vibrational relaxation of CN(v=1) by He. Modeling of the chemistry of circumstellar envelope performed in the laboratory of Rennes included CN+hydrocarbons reactions done in Ar bath. Thus, in order to interpret the

TOMSK STATE UNIVERSITY

## METHODS



Tensor opacities

[5]:

Rate coefficients are obtained by overaging appropriate cross sections over a Bolzman distribution of velocities at a given kinetic temperature *T*:  $(8\beta^3)^{1/2}$   $\int_{-\infty}^{\infty}$ 

$$k_{v \to v'} = \left(\frac{8\beta^3}{\pi\mu}\right)^{1/2} \int_0^\infty \sigma(v \to v')(E_k)$$
$$\times E_k \exp\left(-\beta E_k\right) dE_k$$

Interaction potential is expanded as follows:

level with PBE0 functional in aVTZ basis set plus

(3s 3p 2d 2f) bond functions using SAPT2008

package [2] (SAPT.os for open shell systems).

 $V(r, R, \theta) = \sum_{n=1}^{N_{max}} \sum_{l=1}^{L_{max}} d_{m,0}^{l+m-1}(\cos\theta) A_{ln}(R)(r-r_e)^{n-1}$ 

 $d_{m,0}^{l+m-1}(\cos\theta)$  -reduced Wigner rotation matrix elements, Nmax - number of bond distances

 $k_{NjF\to N'j'F'}(T) = \left(\frac{8}{\pi\mu k_{\rm B}^3 T^3}\right)^{1/2} \int_0^\infty \sigma_{NjF\to N'j'F'} E_{\rm c} \,{\rm e}^{-E_{\rm c}/k_{\rm B}T} \,{\rm d}E_{\rm c}$ 

Thermal rate coefficients at temperature T (average over the collision energy)

 $P^{K}(j \to j') = \frac{1}{2K+1} \sum_{m} |T^{K}(jl; j'l')|^{2}.$ 

 $\sigma_{NjF \to N'j'F'} = \frac{\pi}{k_{NjF}^2} (2F'+1) \sum_{K} \left\{ \begin{array}{cc} j & j' & K \\ F' & F & I \end{array} \right\}^2 P^K(j \to j').$ 

Total energy grid : 0.2 - 1000 cm<sup>-1</sup>  $\rightarrow$  rate coefficients for up to 100 K. The integration range was set from 4.5 a<sub>0</sub> to 80 a<sub>0</sub>. Scattering calculations were carried out using HIBRIDON [6]  $\beta = (k_B T)^{-1} \qquad \qquad E = E_k + \epsilon_v$ 

E – total energy, Ek – kinetic energy,  $\varepsilon v$  – energy of vibrational level

Total energies: up to 15 000 cm-1  $\rightarrow$  converged rate coefficients up to 1500 K.

### RESULTS





Strong propensity rule for Δj = ΔN transitions (predicted theoretically [9] for collisions of molecules in <sup>2S+1</sup>Σ electronic state).
 Strong propensity for transitions with even ΔN.

Near-homonuclear symmetry of the PES strongly favors transitions with even  $\Delta N$  (as in Ref. [10], for example).



Fig.2 Temperature variation of the hyperfine-resolved CN–H<sub>2</sub> rate coefficients for N = 2, j,  $F \rightarrow N' = 1$ , j', F' transitions. Panel (a) corresponds to  $\Delta j = \Delta N$  transitions. Panel (b) corresponds to  $\Delta j \neq \Delta N$  transitions.

 $\Delta j = \Delta N$  transitions (Fig. 2a): strong propensity rule in favor of  $\Delta j = \Delta F$  transitions, especially when N increases. This trend is usual for open-shell molecules (for instance, CN-He collisions [11]).

 $\Delta j \neq \Delta N$  transitions (Fig. 2b): no clear propensity rules can be extracted. Two rules: •rate coefficients show propensity in favor of  $\Delta j = \Delta F$ 

transitions

•final distribution seems to be proportional to the degeneracy (2F' + 1) of the final hyperfine level

The complete set of de-excitation rate coefficients with  $N,N' \le 12$  is available online at the LAMBDA [12] website.



### **CN-He/Ar**

For VCC-IOS method matrix elements of the potential between the vibrational states of the CN molecule are required for fixed values of scattering angle  $\theta$  and all the *R*-values.

Elastic process clearly dominates the inelastic one for any T (Fig.5) and the vibrational relaxation of CN(v = 1) by He is much faster than the vibrational relaxation of CN(v = 1) by Ar.

**IRC+10216**: for average temperature and density [13] and He abundance 0.2 of H<sub>2</sub>,  $\tau \sim 100$  s (for

 $V_{v,v'}(R,\theta) = \langle v(r) \mid V(r,R,\theta) \mid v'(r) \rangle$ 

The coupling of v=0 and v=1 levels increases as the scattering coordinate *R* decreases (Fig. 5).



Fig. 5 Cross sections and rate coefficients for CN-He/Ar.

relaxation with He) in the hottest and densest region. Even in this zone,  $\tau$  is still a minimum of two orders of magnitude higher than the average time  $CN(v=1) + He \longrightarrow CN(v=0) + He$ 

between collisions. Thus, the relaxation by He doesn't constitute a significant competitive channel.

The main collisional partner in curcumstellar envelopes is H<sub>2</sub>. We can estimate, that rate coefficients for  $CN(v=1)+H_2 \rightarrow CN(v=0)+H_2$  will be larger then for  $CN(v=1)+He \rightarrow CN(v=0)+He$  (see, [14]) and, thus, may become a more competitive channel for the destruction of CN (v=1) (for T< 300 K).

Hovewer, there are also reactive collisions of CN with  $H_2$ : CN+ $H_2$  $\rightarrow$ HCN+H. At high temperatures (>1000 K) this pathway may be the dominant one leading to the destruction of CN(v=1). This will be consistent with the fact that CN has not been observed in excited vibrational state in IRC+10216.



with He.

1E4 1E6 1E8 1E10 1E12 1E14 1E16 1E18 1E19

[1] Molpro 2010 http://www.molpro.net

[2] SAPT2008 <u>http://www.physics.udel.edu/~szalewic/SAPT/SAPT.html</u>
[3] Alexander M. JCP, 76, 3637 (1982).
[4] Alexander M. H., Dagdigian P. J. JCP, 83, 2191 (1985).
[5] ] Smith I.W. M., 1980, Kinetics and Dynamics of Elementary Gas Reactions. Butterworths, London

#### References

[6] HIBRIDON <u>http://www2.chem.umd.edu/groups/alexander/</u>
[7] Parker G.A. and Pack R.T. JCP, 68, 1585 (1978)
[8] MOLSCAT V14. Hutson J.M. And Green S. (1994).
[9] Alexander M., Smedley J. E., Corey G. C., 1986, JCP, 84, 3049.
[10] McCurdy C. W., Miller W. H., 1977, JCP, 67, 2488.

[11] Lique F., Kłos J., 2011, MNRAS, 413, L20.
[12] http://www.strw.leidenuniv.nl/~moldata
[13] Agundez M. et al. A&A, 543, 48 (2012).
[14] Lique F. et al. A&A, 478, 567 (2008).

#### **Aknowledgements**

This research was supported by the CNRS national program "Physique et Chimie du Milieu Interstellaire". Y.K. and F.L. are grateful for the financial support of the the CPER Haute-Normandie/CNRT/"Energie, Electronique, Matériaux".