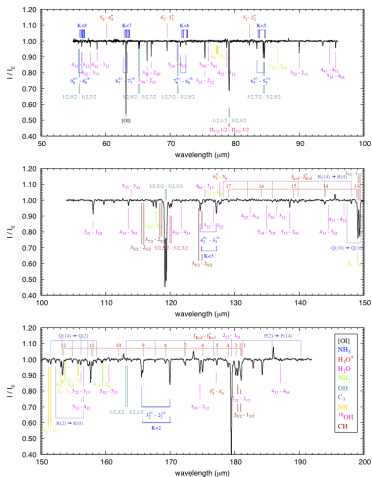


A complete model of  $\text{CH}^+$  rotational excitation  
including radiative and chemical pumping processes

B. Godard & J. Cernicharo

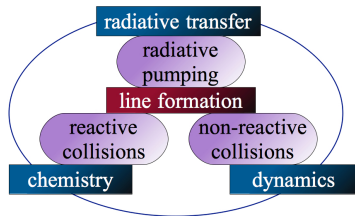
Centro de Astrobiología, INTA-CSIC, Madrid



Herschel/PACS - SgrB2(M)

find minimal nb of components

- density
- kinetic temperature
- velocity field
- abundance

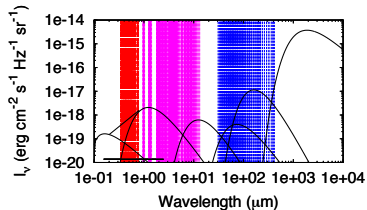
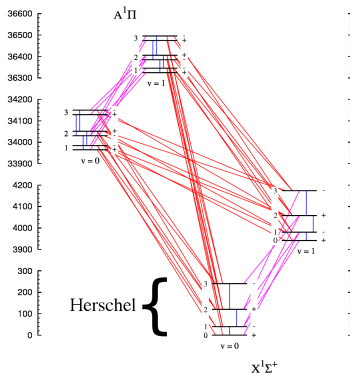


	Orion Bar		NGC 7027	
	$n_{\text{H}}$	$T_{\text{K}}$	$n_{\text{H}}$	$T_{\text{K}}$
OI ( $^3\text{P}_1$ - $^3\text{P}_2$ ), CI ( $^3\text{P}_1$ - $^3\text{P}_0$ ), C <sup>+</sup> ( $^2\text{P}_{3/2}$ - $^2\text{P}_{1/2}$ ), H <sub>2</sub> S(1)	$5 \times 10^4$		$2 \times 10^5$	
CO ( $J < 9$ )	$10^5$	85	$2 \times 10^5$	$\leq 100$
<sup>13</sup> CO ( $J \geq 9$ ), CO ( $J \geq 11$ )	$10^7$	120	$2 \times 10^6$	200
OH ( $^2\Pi_{1/2}$ , $^2\Pi_{3/2}$ )	$10^7$	200	$2 \times 10^6$	1000
CH <sup>+</sup> ( $J \leq 6$ )				
CO <sup>+</sup> ( $2_{5/2} - 1_{3/2}$ )	$\geq 10^6$		$\geq 10^6$	

interclump medium + large density gradients/clumps

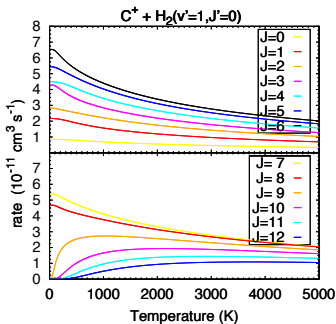
## MADEX - LVG model

- spherical/plan-parallel cloud
- external radiation field
  - ▶ default : local ISRF
  - ▶ scaling :  $T, \beta, \tau, \chi$

energy structure of CH<sup>+</sup>

## formation

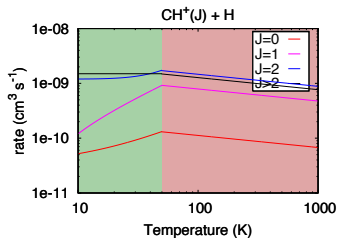
- $C^+ + H_2(v', J') \rightarrow CH^+(v, J) + H$
- others :  $\propto (2J + 1)\exp(-E_J/T_K)$



Zanchet et al. (in prep.)

## destruction

- $CH^+(J) + H$
- others :  $J$ -independent



Plasil et al. (2011)

**parameter domain**

$$[\text{H}_2] = 0.5$$

$$[e^-] = 10^{-4}$$

$$n_{\text{H}} = 10^2 \rightarrow 10^6 \text{ cm}^{-3}$$

$$T_{\text{K}} = 10 \rightarrow 5000 \text{ K}$$

$$\chi_{\text{fir}} = 10^3 \rightarrow 10^8$$

$$\chi_{\text{nir}} = 10^5 \rightarrow 10^{10}$$

$$\chi_{\text{opt}} = 10^2 \rightarrow 10^7$$

main excitation pathways

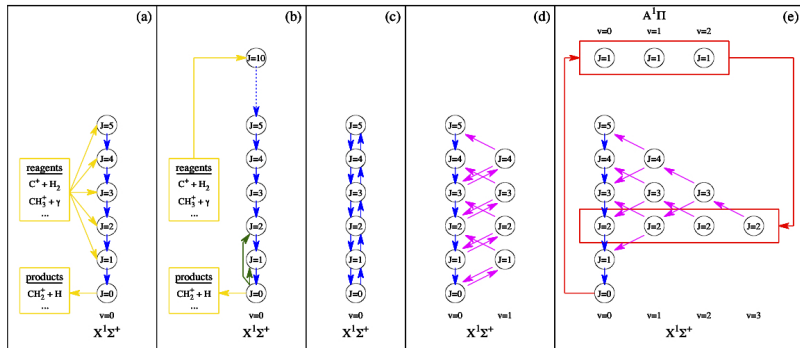
$T_K = 100$

$T_K = 1000$

$\chi_{\text{fir}} = 10^6$

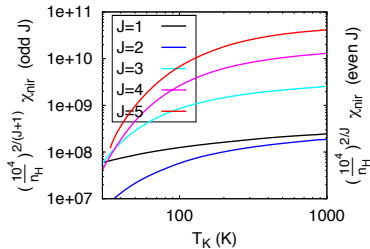
$\chi_{\text{nir}} = 10^{10}$

$\chi_{\text{opt}} = 10^6$



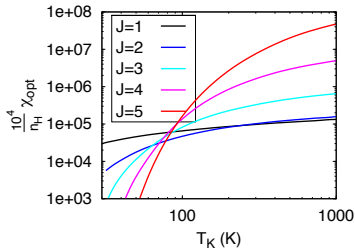
## requirements to activate the radiative pumping

### near-infrared pumping



$$\chi \propto [n_H \exp(-E_J/T_K)]^{2/J}$$

### optical pumping

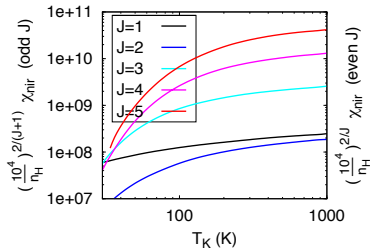


$$\chi \propto n_H \exp(-E_J/T_K)$$



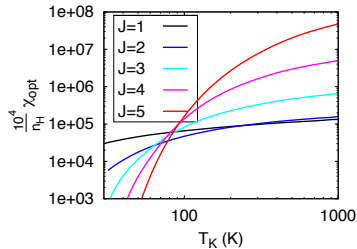
## requirements to activate the radiative pumping

### near-infrared pumping



$$\chi \propto [n_H \exp(-E_J/T_K)]^{2/J}$$

### optical pumping



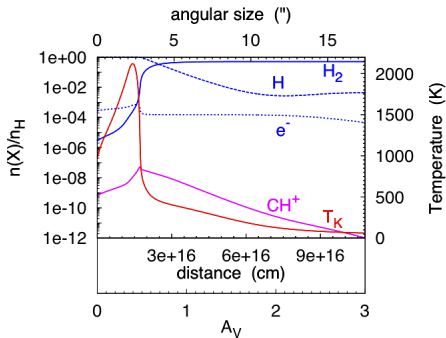
$$\chi \propto n_H \exp(-E_J/T_K)$$

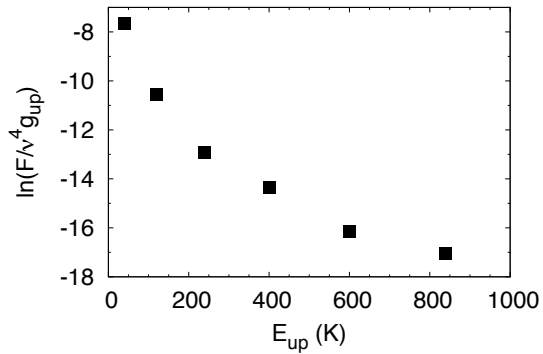
no interstellar clouds with these conditions

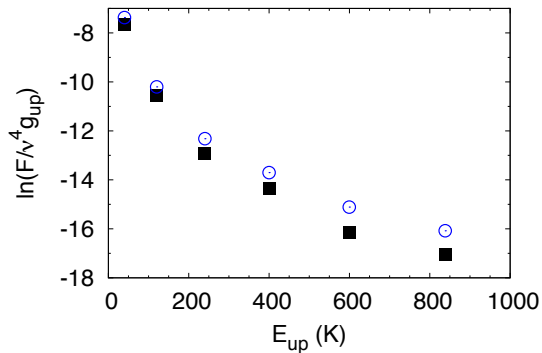
## physical conditions

- $n_{\text{H}} = 5 \times 10^4 \text{ cm}^{-3}$
- $\chi_{\text{nir}} = 3 \times 10^4$
- $\chi_{\text{opt}} = 3 \times 10^4$

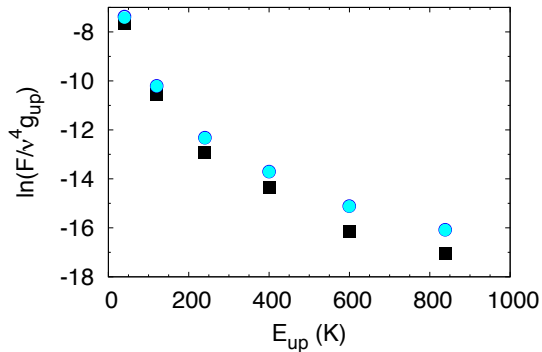
## chemical profiles



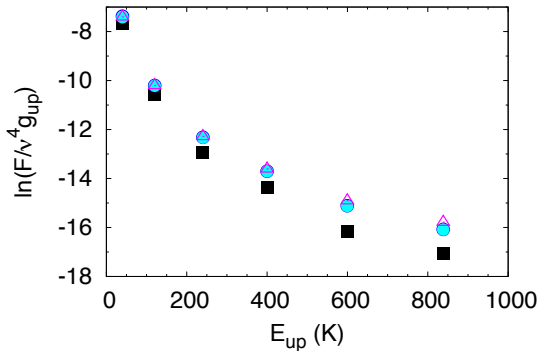




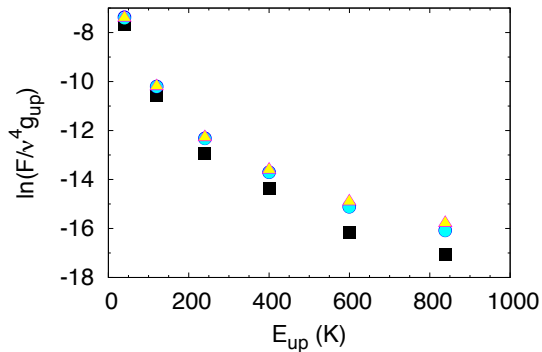
- chemical pumping dominant :  $J \geq 3 - 2$



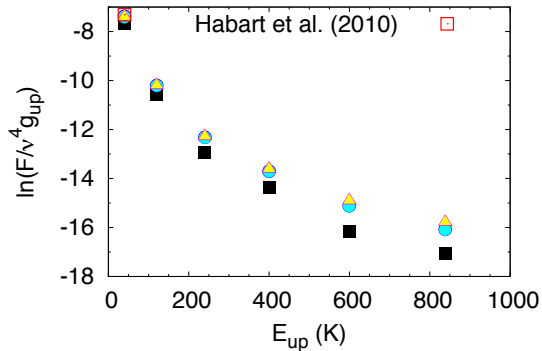
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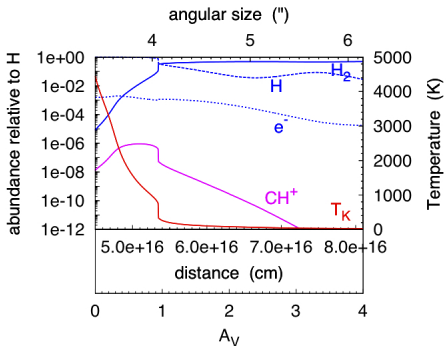
- chemical pumping dominant :  $J \geq 3 - 2$
- Zanchet's rates :  $> 30\%$  on  $J \geq 5 - 4$
- radiative pumping inefficient

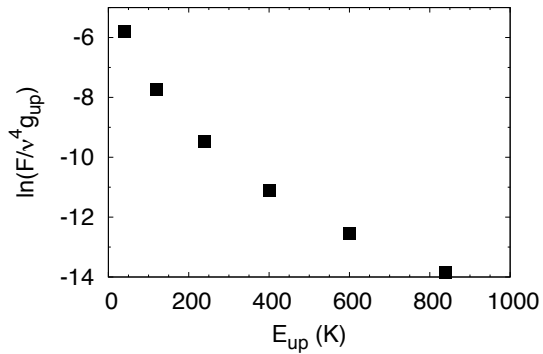


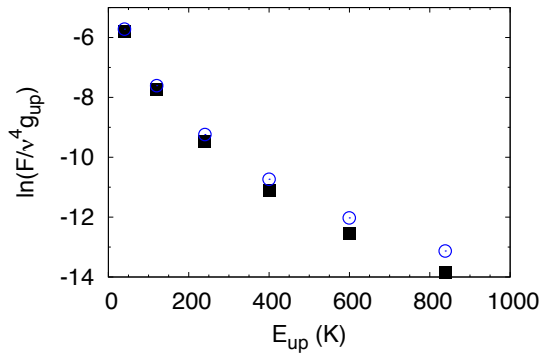
## physical conditions

- $n_{\text{H}} = 2 \times 10^5 \text{ cm}^{-3}$
- $\chi_{\text{nir}} = 2 \times 10^5$
- $\chi_{\text{opt}} = 4 \times 10^4$
- carbon rich  
 $[\text{C}] = 1.3 \times 10^{-3}$

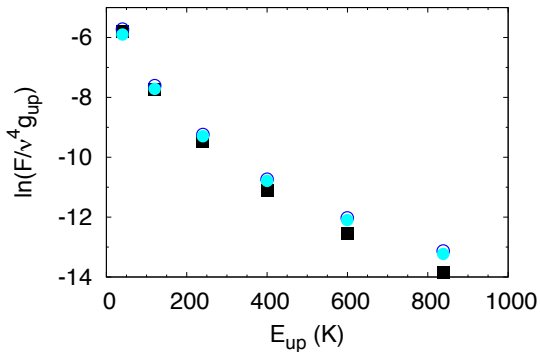
## chemical profiles



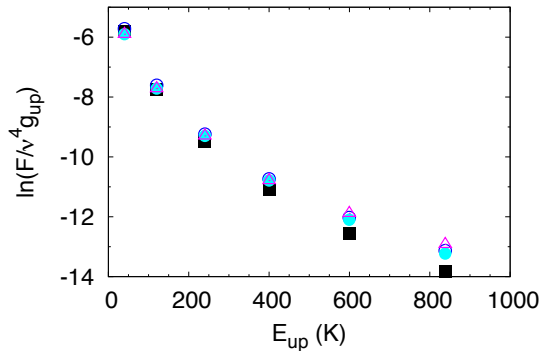




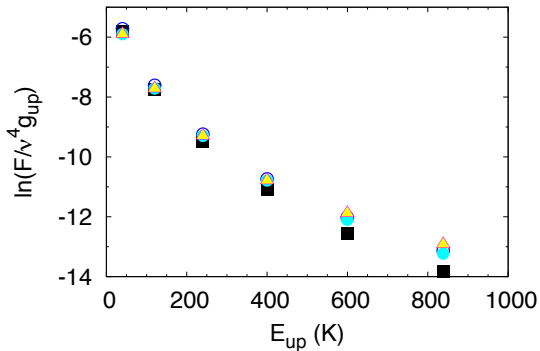
- chemical pumping dominant :  $J \geq 5 - 4$
- collision with  $e^-$  :  $J \leq 4 - 3$



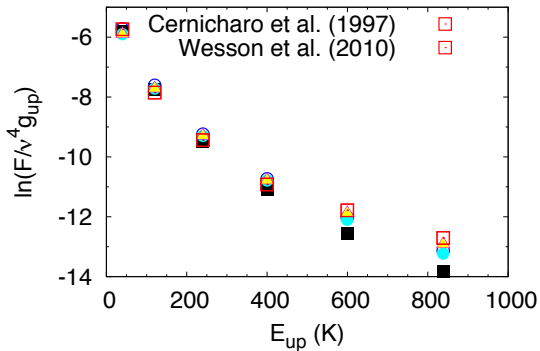
- chemical pumping dominant :  $J \geq 5 - 4$
- collision with  $e^-$  :  $J \leq 4 - 3$



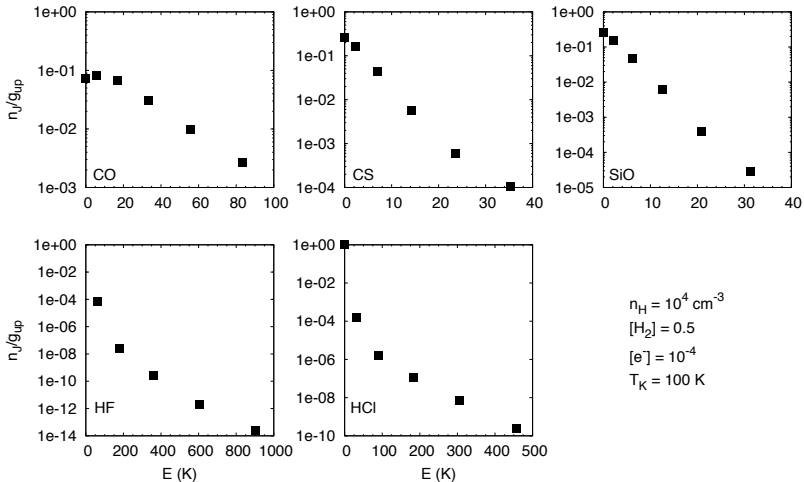
- chemical pumping dominant :  $J \geq 5 - 4$
- collision with  $e^-$  :  $J \leq 4 - 3$



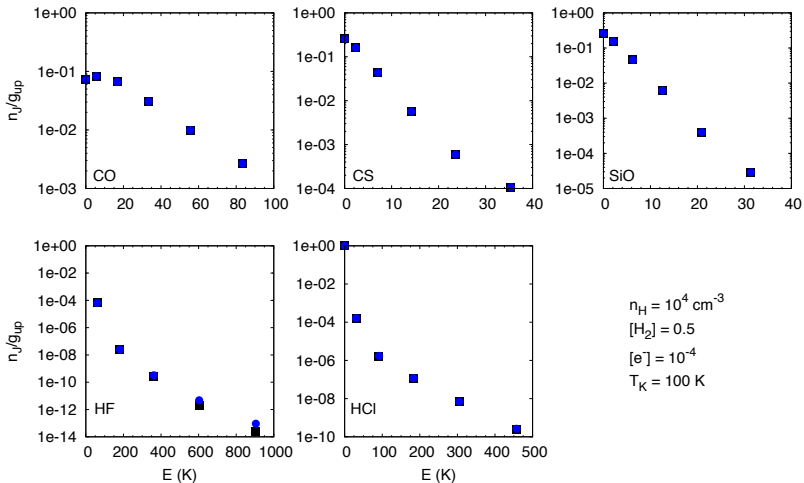
- chemical pumping dominant :  $J \geq 5 - 4$
- collision with  $e^-$  :  $J \leq 4 - 3$
- radiative pumping inefficient

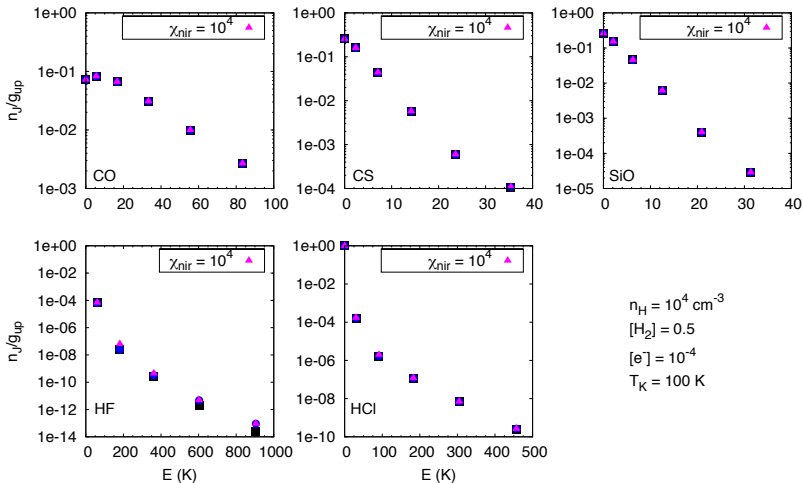


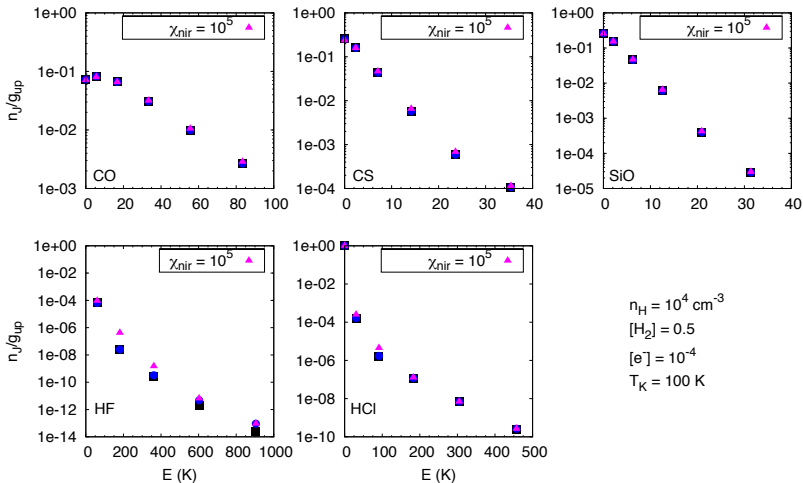
- chemical pumping dominant :  $J \geq 5 - 4$
- collision with  $e^-$  :  $J \leq 4 - 3$
- radiative pumping inefficient
- density 10 times smaller than previous models

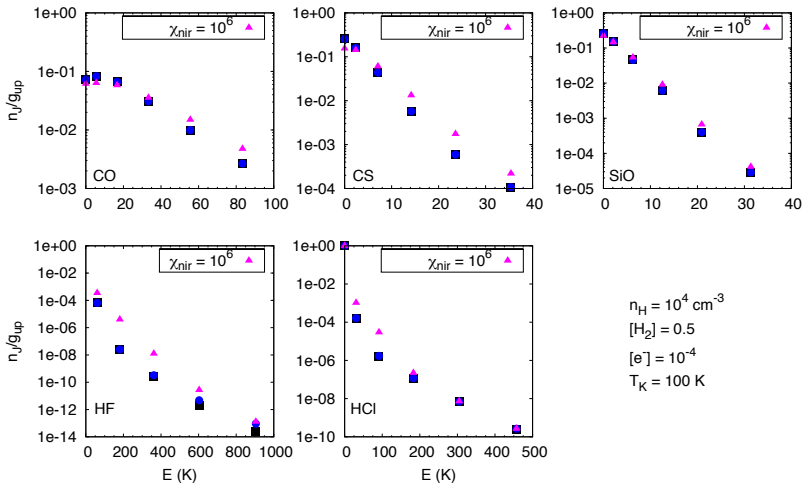


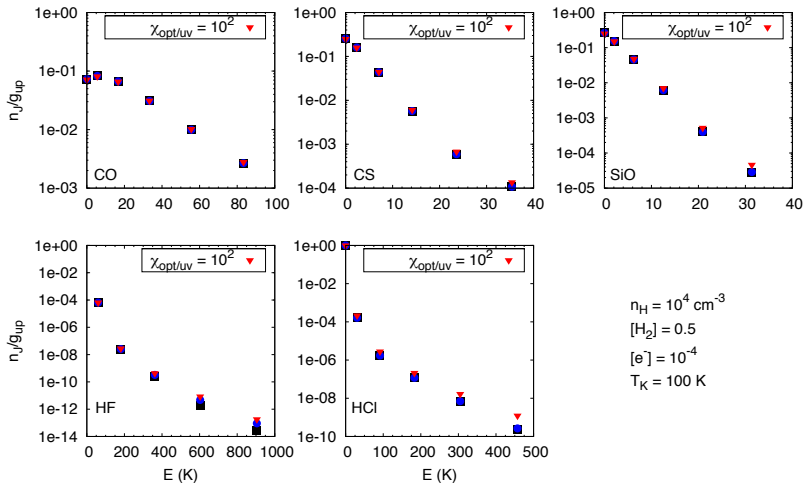


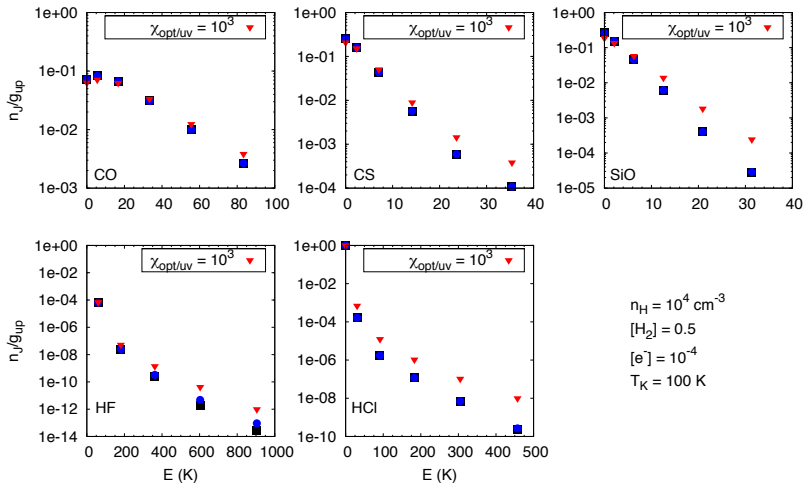


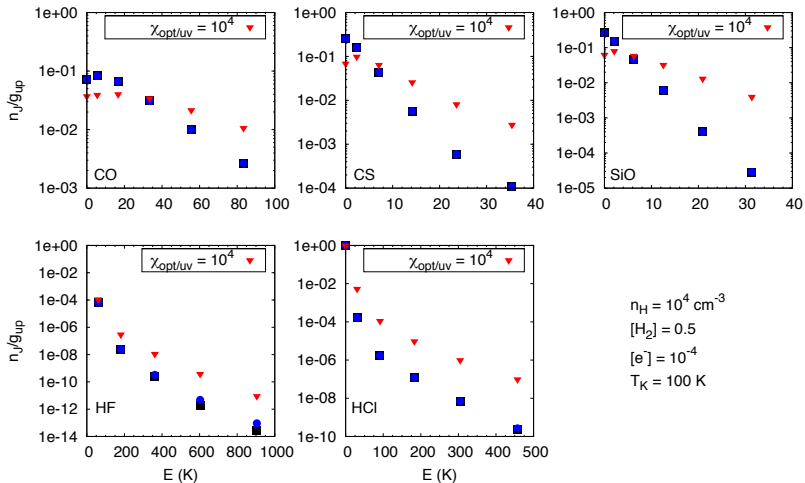












## Summary

- species sensitive to chemical pumping
  - ▶ short lived molecules : e.g.  $\text{CH}^+$ ,  $\text{CO}^+$  (?), ...
  - ▶ state-specific rate constants are important
  - ▶  $\text{CH}^+$  rotational lines explained in low density medium
- species sensitive to near-infrared pumping
  - ▶ HF, HCl for  $\chi_{\text{nir}} \geq 10^4 \times (n_{\text{H}}/10^4)^{2/J}$
- species sensitive to optical/uv pumping and photodissociation
  - ▶ SiO, HCl, CS, and CO for  $\chi_{\text{opt/uv}} \geq 10^3$

## Open issues - Future contributions

- extension to ...
  - ▶ polyatomic molecules : metastable lines of  $\text{NH}_3$
  - ▶ other ground-state configurations : OH ( $^2\Pi$ ) and  $\text{CO}^+$  ( $^2\Sigma^+$ )